

MgB₂ HEB heterodyne receivers at 20-30 K for S & M-sized FIR space missions

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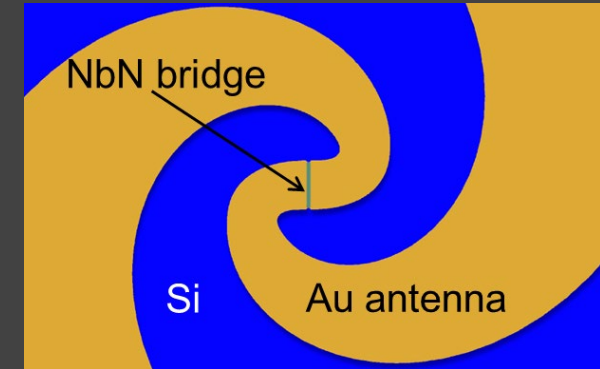
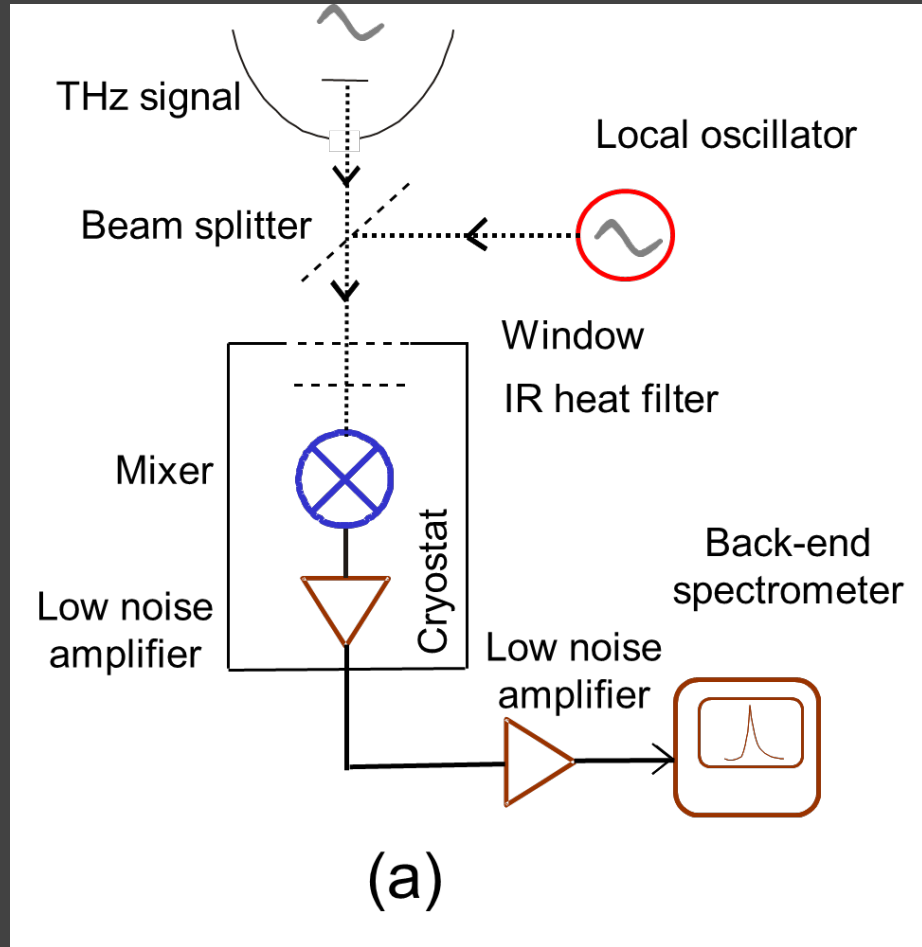
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HEB heterodyne receiver

High spectral resolution ($\nu/\Delta\nu \geq 10^6$)



- NbN HEB mixer
- 1.4 – 5.3 THz
- Operated at 4 K
- Low T_{mixer}^{DSB} (optics/LNA+ $T_{mixer}^{DSB} \rightarrow T_{rec}^{DSB}$)

THz space observatories with NbN HEB mixers

Herschel



HIFI: 1.4 – 1.9 THz

ESA

Cherednichenko, et al, Rev. Sci. Instrum., (2008)

SRON

Stratospheric Observatory
for Infrared
Astronomy(SOFIA)



1.83 – 2.06 THz, 4.7 THz

C. Risacher et al.,
IEEE Trans. Terahertz Sci. Technol., (2016)

The Stratospheric Terahertz
Observatory (STO2)



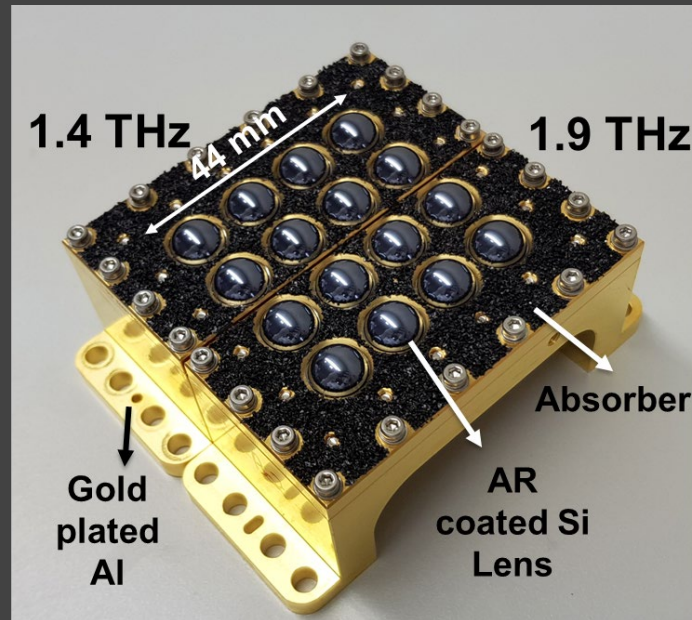
1.4 THz, 1.9 THz, 4.7 THz

GUSTO with NbN HEB arrays

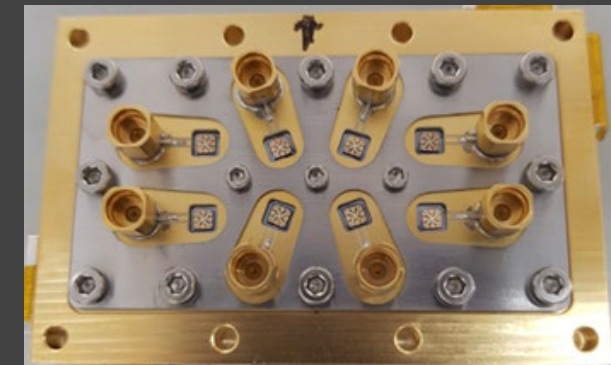
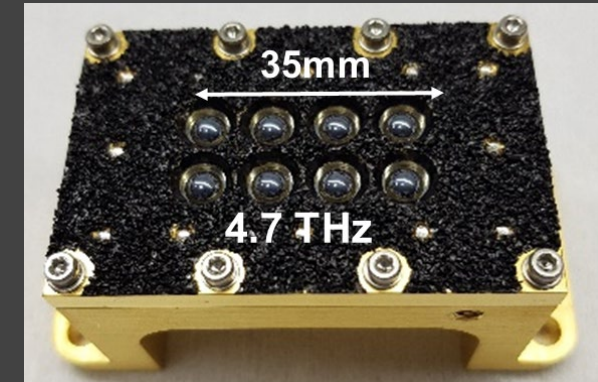


C. Walker et al, Proc. SPIE 12190 (2022);

- NASA Explore mission with 3 HEB arrays on board
- Flown above Antarctica for 57 days
- Mapping [NII] , [CII] lines for 62 square degrees + LMC (see poster by Russ Shipman)



J. R. G. Silva et al ,
J. Astron. Telesc. Instrum. Syst. (2025)

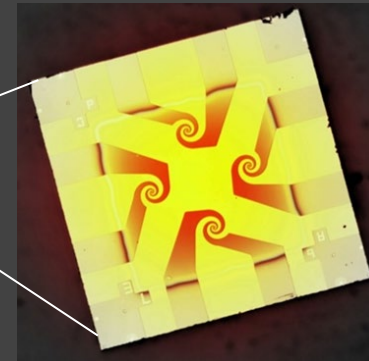


More sensitive NbN HEB mixers available

LO frequency	$T_{\text{mixer}}^{\text{DSB}}$	$L_{\text{mixer}}^{\text{DSB}}$	$T_{\text{mixer}}^{\text{DSB}}$ (Quantum noise units)
1.63 THz	$240 \pm 6 \text{ K}$	$4.4 \pm 0.09 \text{ dB}$	$3.1 \times h\nu/k_B$
2.52 THz	$290 \pm 13 \text{ K}$	$5.5 \pm 0.2 \text{ dB}$	$2.4 \times h\nu/k_B$
5.25 THz	$620 \pm 54 \text{ K}$	$7.4 \pm 0.4 \text{ dB}$	$2.5 \times h\nu/k_B$

$T_{\text{mixer}}^{\text{DSB}}$ is in units of $h\nu/k_B$ (QN). Note that all values are obtained at an IF of 1.7 GHz and an operating temperature of 4.6 K.

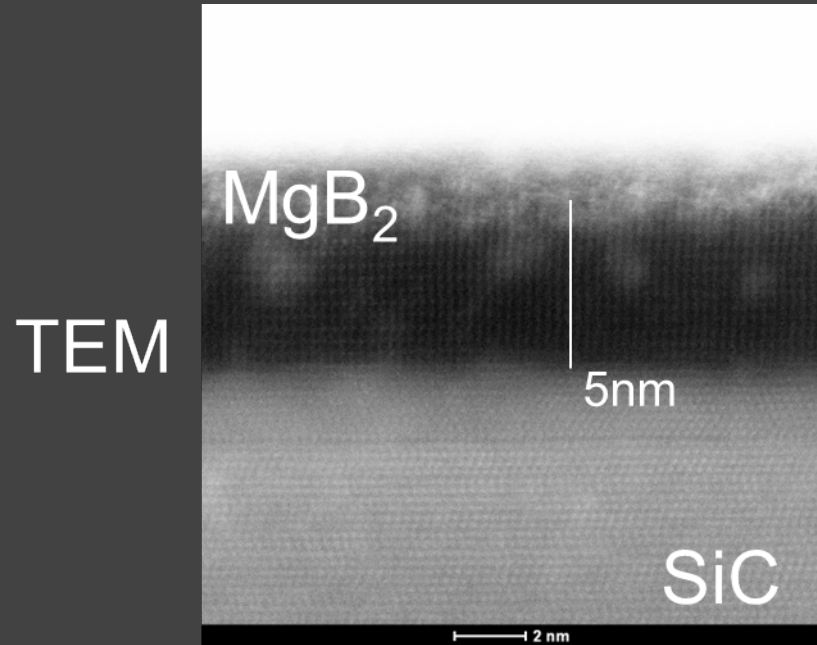
	1	2	3	4	5	6	7	8
A								
B		T					T	
C								
D								
E								
F								
G		T					T	
H								



- Realized even more sensitive NbN HEBs for OASIS and SALTUS
- More than 30 % improvement in $T_{\text{mixer}}^{\text{DSB}}$ than GUSTO HEBs (1.6 and 2.5 THz)
- 60 chips ready for future missions, e.g. LETO for M8

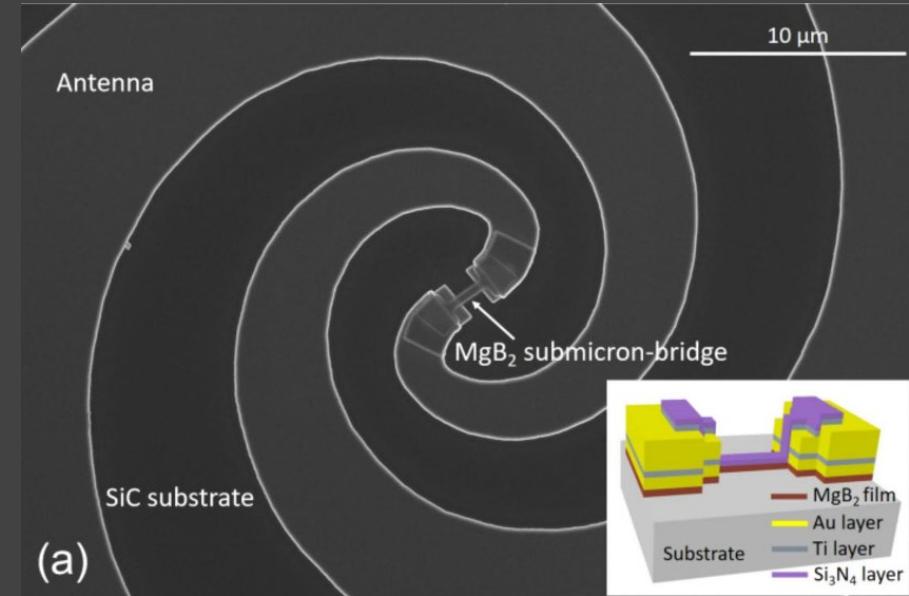
B. Mirzaei, J. R. G. Silva, W. -J. Vreeling, W. M. Laauwen, D. Ren and J. -R. Gao,
IEEE Trans. Terahertz Sci. Technol., (2025)

Alternative technology: MgB_2 HEBs



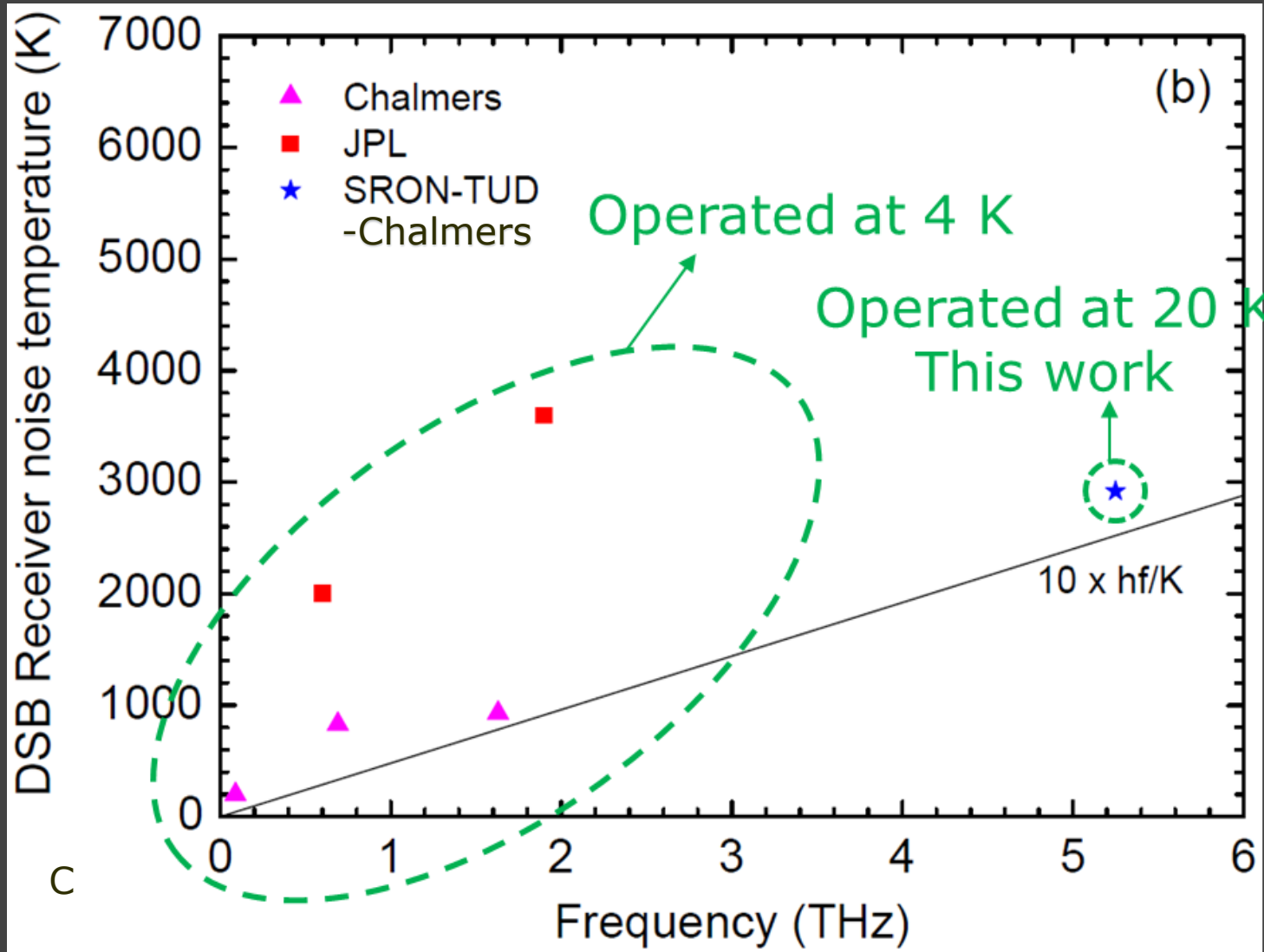
Magnesium diboride (MgB_2) films grown by Hybrid Physical CVD at Chalmers

- MgB_2 films (6-8 nm) with a T_c of 34–40 K → High operating temperature of ≥ 20 K



S. Cherednichenko et al, Supercond et al, Sci. Technol. 34 044001 (2021).

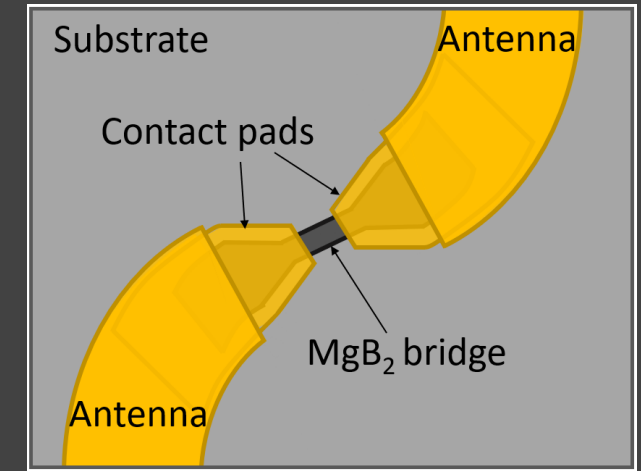
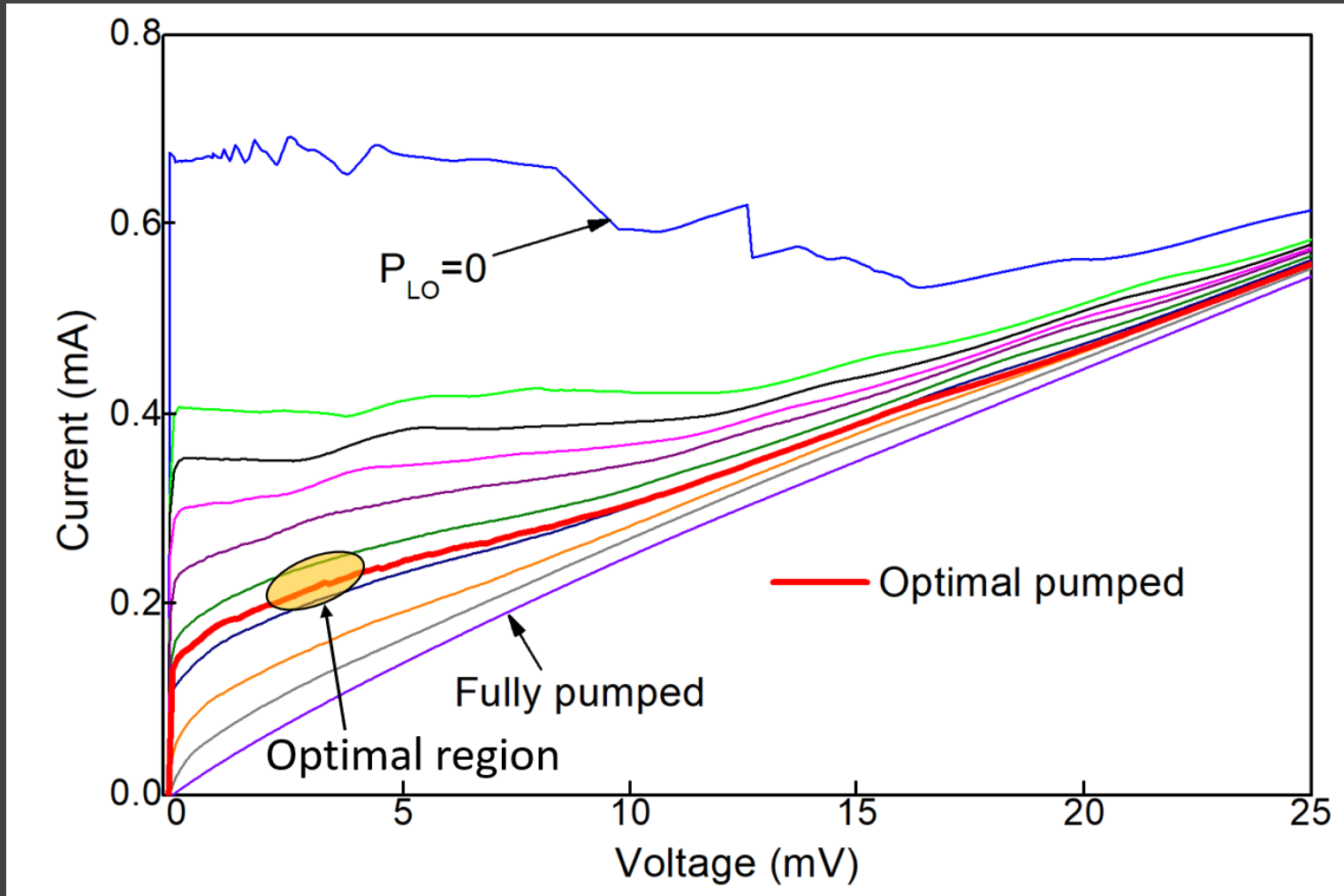
MgB₂ HEB mixers: results so far



IF bandwidth of ≥ 10 GHz

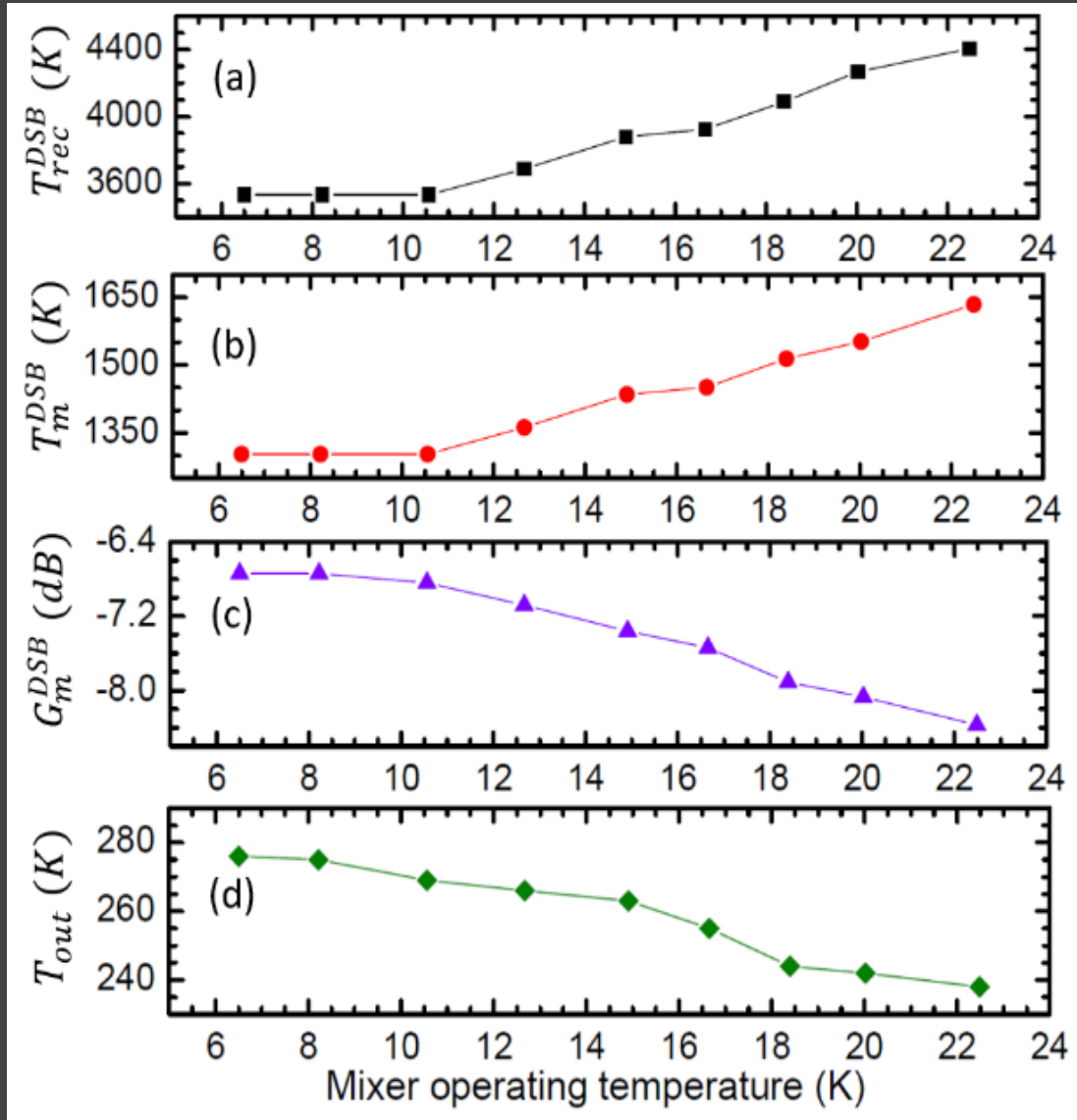
- Novoselov, E. and Cherednichenko, S., Appl. Phys. Lett. (2017). Also reports at 15K and 20K.
- Cunnane, D., Kawamura, J. H., Wolak, M. A., Acharya, N., Tan, T., Xi, X. X., Karasik, B. S., Appl. Supercond.(2017).
- Gan, Y., Mirzaei, B., Silva, J. R. G. D., Chang, J., Cherednichenko, S., van der Tak, F., and Gao, J. R. Appl. Phys. Lett. (2021).
- Y. Gan, B. Mirzaei, J. R. G. D. Silva, S. Cherednichenko, F. van der Tak, and J. R. Gao, J. of Appl. Phys. (2023)

IV curves at 5.3 THz



- $T_c=38.4$ K, $W=280$ nm, $L=850$ nm, $R_N=51$ Ω ,
- Optimal pumped IV gives LO power of 10 μ W

Measured temperature dependence at 5.3 THz



$$T_{rec}^{DSB} = T_{opt} + \frac{T_{mixer}^{DSB}}{G_{opt}} + \frac{T_{IF}}{G_{opt} G_{mixer}^{DSB}}$$

- T_{mixer}^{DSB} increased by 17 % from 6 to 20 K
- T_{mixer}^{DSB} (1540 K) at 5.3 THz is 2.5 times the best NbN HEB mixer
- T_{mixer}^{DSB} (950 K) at 1.63 THz is 4 times the best NbN HEB mixer
- can be improved by cleaning the interface between MgB_2 and Au

MgB₂ HEBs can be operated at a 20 K cooler

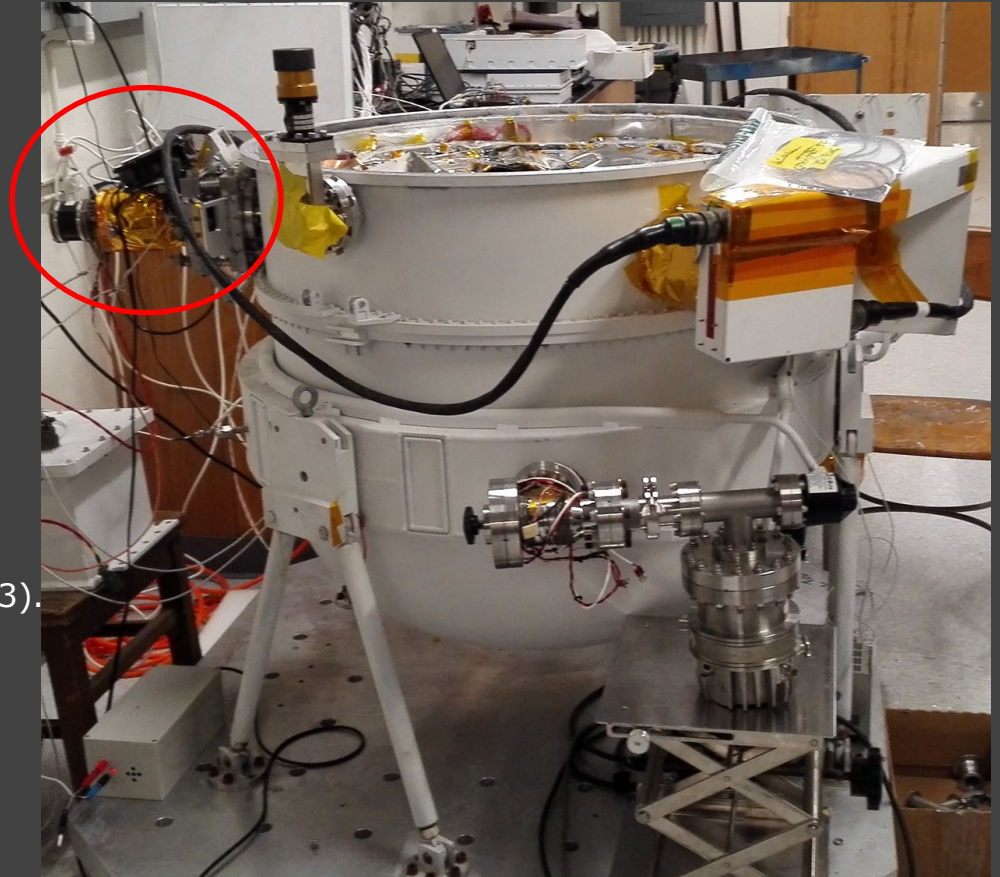
- Apply a 20 K Stirling cooler (XRISM)



Y. Sato et al, J. of Space Technology and Science, (2013).

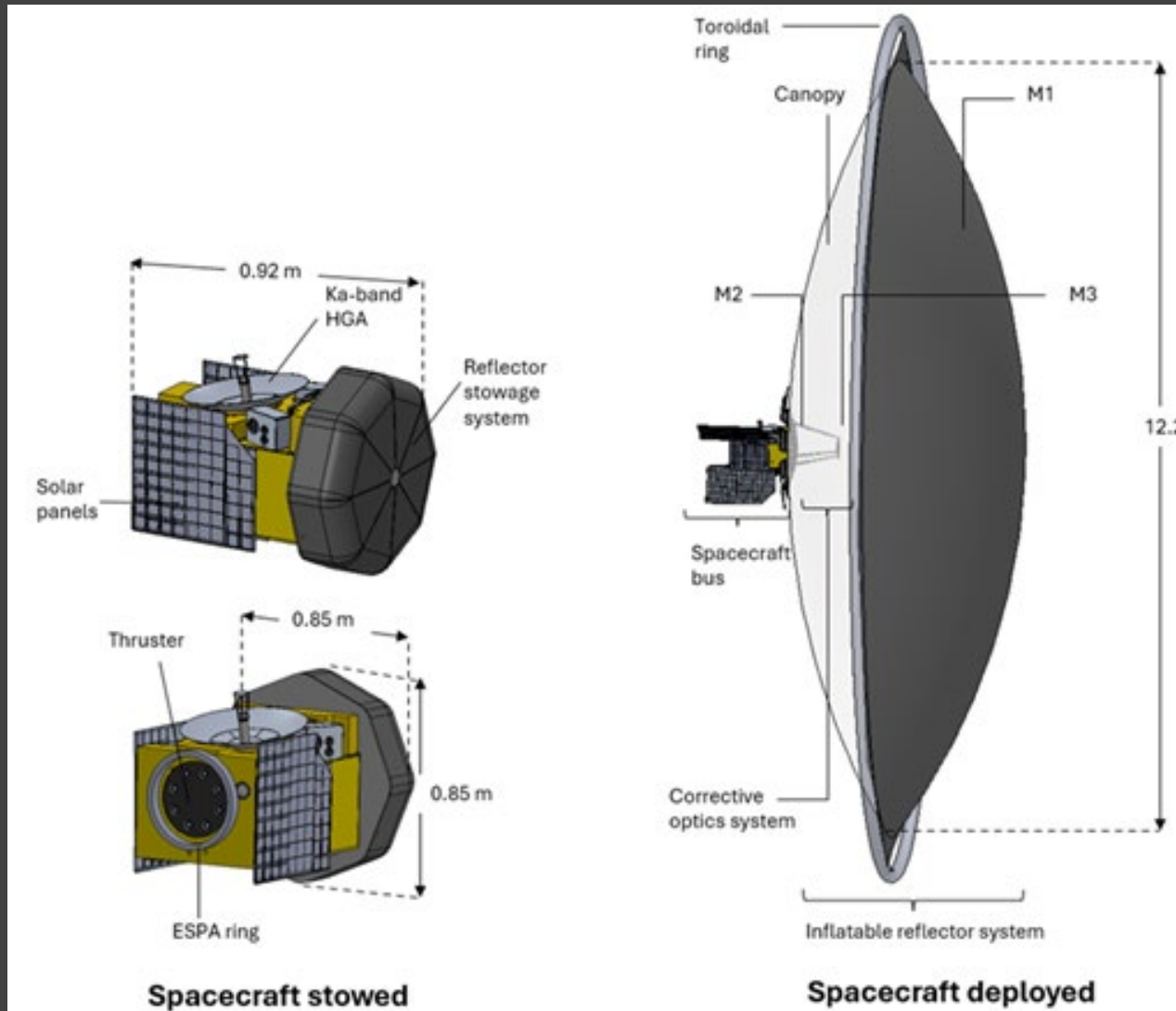
- MgB₂ HEBs enable S & M FIR missions
- MgB₂ HEBs facilitate Space FIR Interferometers

STO2 4K L-He cryostat



MgB₂ HEBs for TOP, proposed to NASA “Pioneers”

- Terahertz Orbiting Pioneer (TOP)

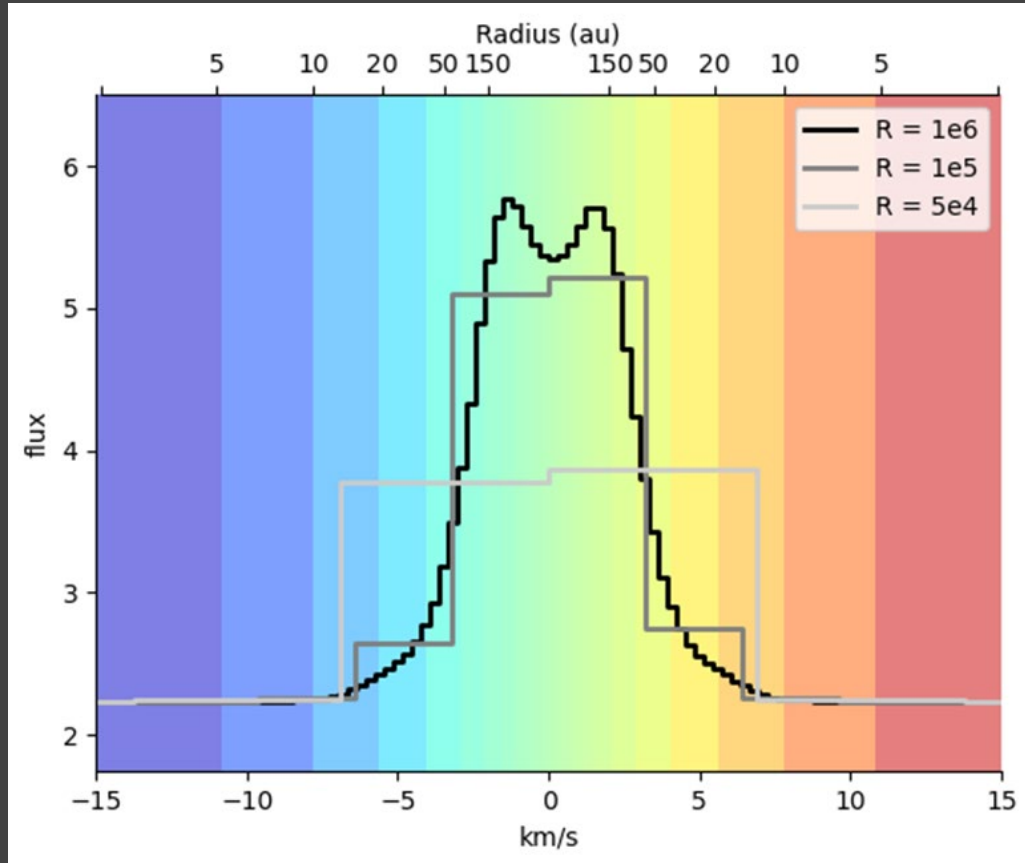


- Inflatable telescope, effective diameter of 8 meter
- Introduce a 20 K cooler
- With 4 MgB₂ HEB mixers
- HD (1-0)@2.7 THz and (2-1)@5.3 THz
- PI: Chris Walker at UA
- SRON/Chalmers : deliver MgB₂ HEBs

Credit: Chris Walker

TOP has unique science capability

- TOP: Spectral resolution $R=10^6$



Credit: Chris Walker and Peter Roelfsema

- the mass distribution and dynamics within protoplanetary disks
- Simulated disk-integrated flux emission spectra of GM Aurigae for the HD 1-0 line
- High spectral & spatial resolution

Key takeaways

- MgB_2 HEBs, with limited efforts, have shown promising results between 0.5 to 5.3 THz
- MgB_2 HEBs, requiring a 20 K cooler, enable small & medium FIR missions
- If TOP can fly, TRL of MgB_2 HEBs will increase