

The FIR-Data Gap and how the SOFIA Data Center (SDC) can help







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Abstract

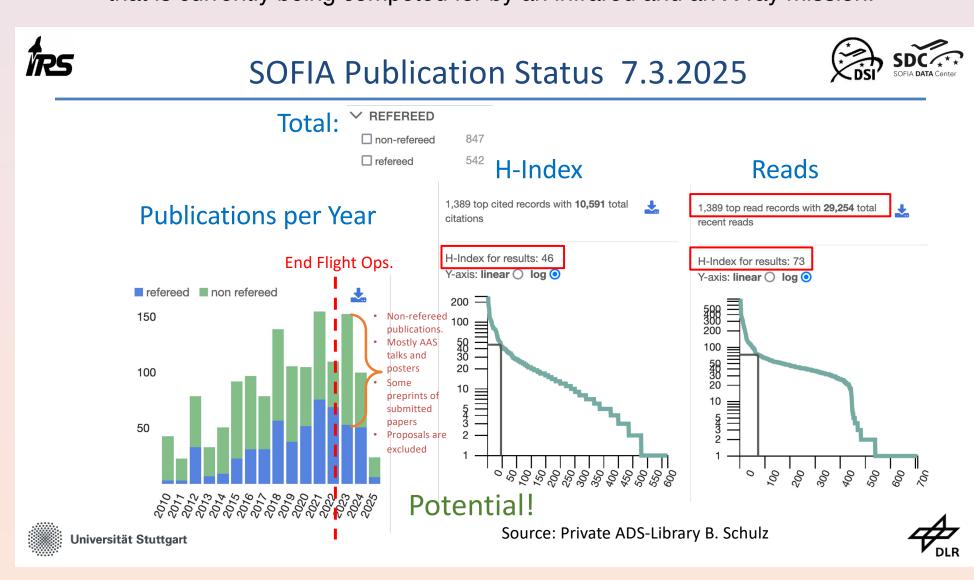
During 783 scientific flights, SOFIA, the Stratospheric Observatory for Infrared Astronomy of the space agencies of Germany and the United States, DLR and NASA, respectively, collected a considerable volume of scientific data, that is available through the Infrared Science Archive at IPAC. After the end of flight operations in September 2022, with only one year of ramp-down time for Science Mission Operations remaining, only a limited data reprocessing of SOFIA Observing Cycles 5 to 9 (2017 to 2022) could be achieved. In order to complete the job and ensure a full exploitation of the considerable investment in SOFIA, the Institute of Spaceflight Systems of the University of Stuttgart, the home of the German SOFIA Institute (DSI), submitted a proposal to DLR to create a SOFIA Data Center (SDC) with a more comprehensive plan for a five-year post operational phase. The proposal was accepted and funded and the project started in July 2024.

This contribution will make a brief assessment of the history and outlook of Far-Infrared (FIR) astronomical observatories, and discuss the SDC's plans to perform a more comprehensive reprocessing of all stored SOFIA data. We will use upgraded data processing pipelines, with improved pointing reconstruction, derived from guide camera imagery and new corrections for atmospheric water vapor and instrumental effects. The science data - re-ordered by astronomical criteria - will be stored along with the technical and operational engineering data of the telescope in a searchable, VO-compatible, publicly accessible archive.

The newly formed SDC draws from the experience of the former DSI and seeks collaboration with scientists and engineers, that still hold knowledge of SOFIA operations and instrumentation. Once the SDC work is complete after 5 years, the archive will be transferred to the new German Astrophysics Center (DZA) as a long-term home. We expect to work with the entire astronomical community to promote the exploitation and publication of valuable Mid- and Far-Infrared data, establish a powerful scientific resource for the post-mission phase, and bridge the gap until the next space- or stratospheric FIR observatory becomes available.

Motivation ÎRS Far-Infrared Astronomy History and Future 57cm Telescope 12μm, 25μm, 60μm, 100μn NASA Ames U2 (early 1970s) Learjet Observatory (late 1960s and 1970s 30.4cm Telescope (replacable) 91cm Telescope (2010-2022) Convair 990 "Galileo" (1965-1973) Universität Stuttgart

With its beginnings in the early 60s, the young field of infrared astronomy went through an exciting time with a number of very successful stratospheric and space-based observatories. However, the premature end of SOFIA, the last regularly scheduled observatory sensitive to the FIR, after less than half of its design lifetime of 20 years, leaves the field with an uncertain future. The earliest realistic chance for an infrared space mission is the NASA Probe mission, that is currently being competed for by an infrared and an X-ray mission.



SOFIA's yearly publication statistics was steadily rising while there were observing flights and support for scientists through the Science Mission Operations Center (SMO). A considerable fraction of the data yet waits for publication. Active work on data improvement and advertising the archive contents among the scientific community will help fully exploit the investment made in this mission.

Aims of the SOFIA Data Center

- Creating a powerful scientific resource for the time after the mission Bridging the expected 10-20 year gap in regular new FIR data availability
- Ensuring the quality of the SOFIA data archive and its documentation to secure its scientific productivity for the long term
- Enabling future scientific data exploitation without the need for expert
- knowledge of SOFIA subtleties Completing the work within 5 years with subsequent transfer to DZA
- Staying in sync with the current SOFIA data repository hosted within

Main Tasks

- Update of metadata to enable reprocessing of Cycles 0 4
- Create an automatic reprocessing environment allowing for an iterative approach
- Reprocess instrument data (GREAT instrument will be separately done in Cologne)
- Iteratively implement additional pipeline improvements
- Develop supporting software tools
- Create a SOFIA Science Archive hosting raw and processed data
- Reorganize science data rather by astrophysical criteria than operational ones
- Create an Operations- and Engineering-Archive
- Update documentation where necessary
- Support scientists through webinars, conferences and direct advice
- Support engineering research by interested parties

Summary

- The SOFIA Data Center (SDC) at the University of Stuttgart will provide an adequate
- post-processing and archival phase for SOFIA.
- The active phase for the SDC is 5 years.
- Within this period the data will be substantially improved, be placed into a well searchable and documented online-archive, and scientists will receive expert support in understanding residual instrumental effects and finding valuable, yet unpublished

Improved Pointing and Atmospheric Water

At the end of this period the archive contents are planned to be transferred into a permanent archive at the DZA.



SOFIA during flight testing with chase plane (Credit: NASA/DLR)

Data Volume in DCS

4846 h shown (Data from 67 Flights ~470h not counted here)

The SOFIA Data Vault

SOFIA's eight science instruments collected a considerable amount of data during her

operational lifetime. The data comprises the instrument science data, instrument and

observatory housekeeping data and all images taken in parallel by the three telescope guider

cameras. After observing flights ceased, there was only limited time and funding to perform

data reprocessing. These activities focused on Cycles 5 - 9 and ended after one year.

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Pfüller, E., Wolf, J., Wiedemann, M. 2018, Journal of Astronomical Instrumentation,

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TRS

• SOFIA Raw Data ~141 TB

Total about 5300 h (SLT)

Telescope guider cameras FPI+, WFI

References

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Volume 7, Issue 4, id. 1840006

475, 315

und FFI were always active (>5300 h).

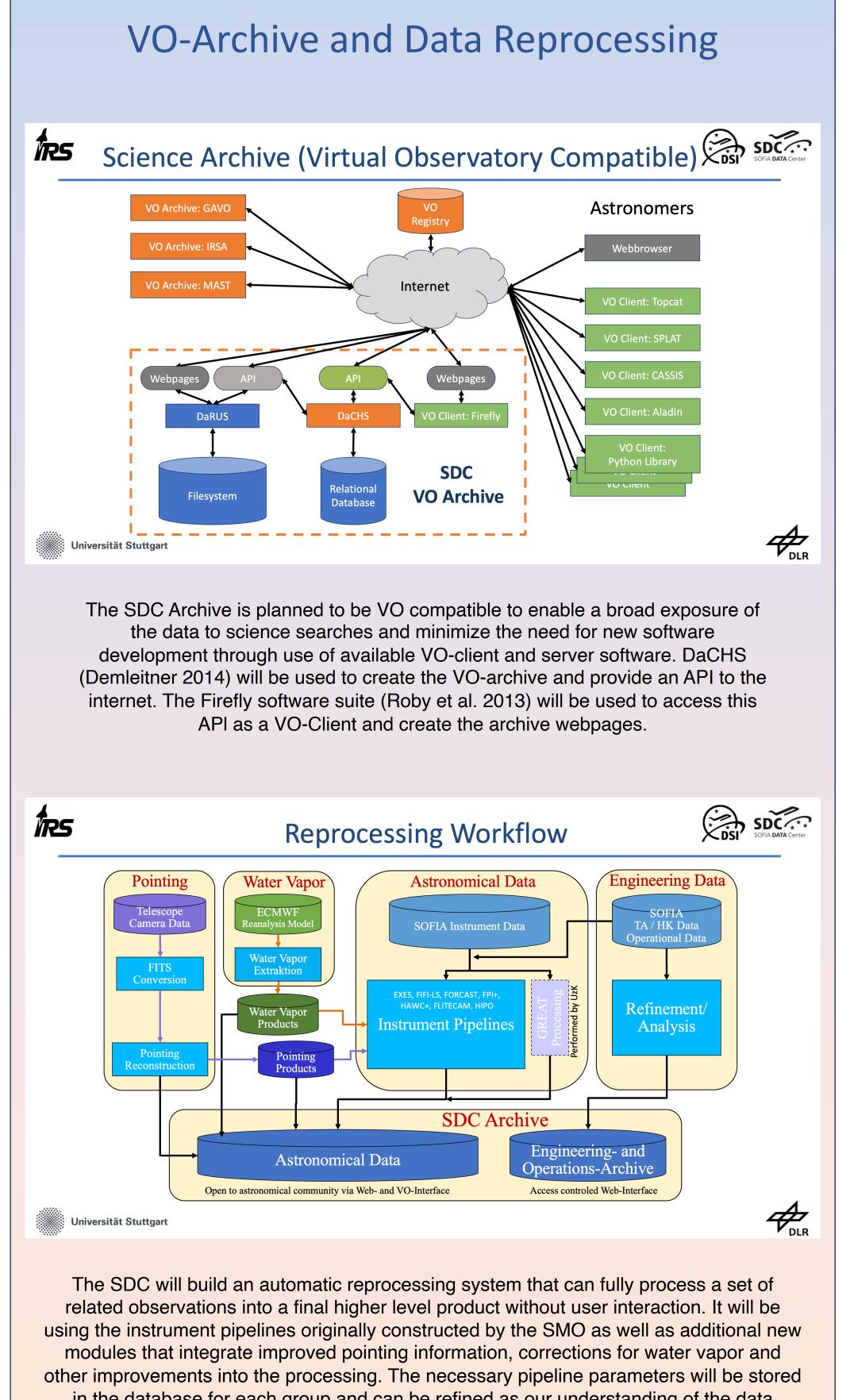
• Science Data ~24 TB

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SDC **
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modules that integrate improved pointing information, corrections for water vapor and in the database for each group and can be refined as our understanding of the data improves. The grouping of related observations will be driven by scientific considerations. Groupings that arose only from operational constraints will be eliminated.



Vapor Correction TA-Camera Images and Pointing Reconstruction SDEC SOFIA DATA Center 9.8 mag CCD Images to be released in FITS-format for serendipity research Telescope boresight pointing reconstruction 11.2 mag via plate solving **Example Coverage** 17-Jan-2020 - 13-Mar-2020 The three CCD cameras WFI, FFI, and FPI+ (Pfüller et al. 2018) of the telescope assembly (TA) recorded images through all science flights. The SDC will provide the images with FITS headers, publicly release them and generate detailed telescope pointing reconstruction products for every flight that can be used by the data processing SDC ** SOFIA DATA Center ÎRS Water Vapor Calibration The FIFI-LS correction uses water **SOFIA Flight Path** vapor measurements before and after measuring the object. Precipitable Water The ECMWF method covers all changes during the observation. **Example Mars Spectrum** In-situ measurements of water-vapor overburden with FIFI-LS correlate well with values derived from ECMWF ERA5 atmospheric re-analysis model. Fischer + 2021, Iserlohe + 2021, Iserlohe + 2022, Vacca + 2023 A new method to calibrate the atmospheric transmission mainly due to its water vapor content is based on a global atmospheric reanalysis provided by the European Centre for Medium-Range Weather Forecasts (ECMWF). The method was validated using dedicated, in-flight FIFI-LS observations. It can provide corrections for all data from all SOFIA flights even after the fact (Fischer et al. 2012, Iserlohe et al. 2021, Iserlohe et al.

2022, Vacca et al. 2023). In particular, data from instruments with low to intermediate

spectral resolution will benefit from an implementation of this method into their

corresponding processing pipelines.

Workshop: 2-4 of April 2025, Leiden, The Netherlands. "Future Role of FIR-Submm Space Observations - Defining Urgent Science and Instrumentation" In honor of Thijs de Graauw and his collaborators

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